

Extension Worksheet – Option E, Worksheet 2, HL only

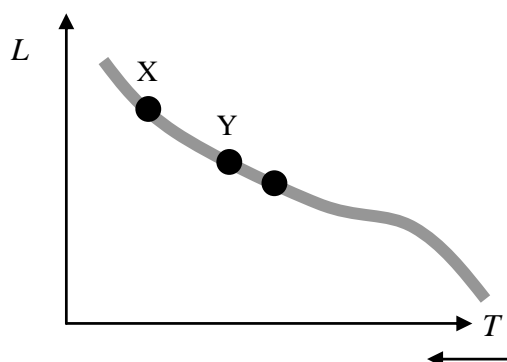
- 1 A large, cool cloud of gas may collapse under gravity to form a star. State where the energy comes from to heat up the star so that nuclear fusion may take place. [1]

- 2 In this question assume a mass–luminosity relation of $L \propto M^{3.5}$.

- a Suggest why the quantity $\frac{L}{M}$ is an estimate the lifetime on the main sequence of a star of mass M and luminosity L . [2]

- b Estimate the lifetime on the main sequence of a star of mass $12 M_{\odot}$ in terms of the Sun's lifetime T_{\odot} . [2]

- 3 The HR diagram below shows two main sequence stars, X and Y, and our Sun.



- a Explain how it may be deduced that star X is the most massive of the three stars. [2]
- b Draw lines on the diagram to show the evolutionary paths of stars X and Y. [2]
- 4 Two main sequence stars have masses of $2 M_{\odot}$ and $10 M_{\odot}$. Compare the evolutionary paths of these two stars with reference to
- a their final fate. [2]
- b the intermediate stages in their evolution. [2]
- c the nuclear reactions taking place in the stars. [3]
- 5 Outline how the following stars manage not to collapse under their own weight.
- a A main sequence star. [1]
- b A white dwarf. [1]
- c A neutron star. [1]
- 6 Neutron stars are not shown on an H-R diagram. Suggest a reason for this. [1]
- 7 State and explain whether the inverse of the Hubble constant underestimates or overestimates the age of the universe. [2]
- 8 Explain why, in the very early universe, it was not possible to have stable structures such as nuclei and atoms. [2]